

1 Continental weathering in the Early Triassic in Himalayan  
2 Tethys, central Nepal: Implications for abrupt environmental  
3 change on the northern margin of Gondwanaland

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5 **Kohki Yoshida<sup>1</sup>, Toshio Kawamura<sup>2</sup>, Shigeyuki Suzuki<sup>3</sup>, Amar Deep Regmi<sup>4</sup>, Babu Ram**  
6 **Gyawali<sup>5</sup>, Yuka Shiga<sup>1</sup>, Yoshiko Adachi<sup>7</sup> and Raj Megh Dhital<sup>4</sup>**

7 *<sup>1</sup>Department of Geology, Faculty of Science, Shinshu University, Matsumoto 390-8621, Japan,*  
8 *kxyoshid@shinshu-u.ac.jp*

9 *<sup>2</sup> Department of Earth Science, Miyagi University of Education, 980-0845, Japan*

10 *<sup>3</sup> Department of Earth Science, Okayama University, Okayama 700-8530, Japan*

11 *<sup>4</sup> Central department of Geology, Tribhuvan University, Kritipur, Kathmandu, Nepal*

12 *<sup>5</sup> Department of Earth Science, Tohoku University, Sendai, 980-8578, Japan*

13 *<sup>6</sup> Chiyoda Corporation, Yokohama, 220-8765, Japan*

14 *<sup>7</sup> Department of Geology, Faculty of Science, Niigata University, Niigata 950-2181, Japan*

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16

17 **ABSTRACT**

18           The geochemistry of Triassic mudstones in the Himalayan Tethys sequence, central  
19 Nepal, was studied with respect to changes in sedimentary facies, grain size, and source rocks.  
20 The Triassic sedimentary facies of mudstone and carbonates show deposition in offshore to  
21 hemiplegic environments. The rare earth element (REE) pattern of the Permian and Triassic  
22 mudstones suggests uniformity correlatable to average shale. The major element geochemistry of  
23 the Early Triassic Griesbachian–early Smithian mudstones indicates a sediment supply from  
24 strongly weathered sources with the chemical index of alteration (CIA) values of 76–81.  
25 However, the mudstones in the late Smithian show weakly weathered sources with CIA values of  
26 68–74. The lower part of the Middle Triassic Anisian mudstones return to Early Triassic  
27 paleoweathering levels. There are no significant relationships among lithofacies, the grain size of  
28 the sediments, and CIA values. Thus, the abrupt change of the degree of paleoweathering in the  
29 Early Triassic, late Smithian time, suggests a dramatic decrease in continental weathering, which  
30 is related to a predominantly arid climate in the northern marginal area of Gondwana.

31

32 **INTRODUCTION**

33           The end-Permian mass extinction in the late Paleozoic era, approximately 250 million  
34 years ago, marked the disappearance of the typical end-Paleozoic faunas, which dramatically  
35 affected the evolution of life (Erwin, 1993; Raup and Sepkoski, 1982; Hallam, 1991). A  
36 significantly delayed recovery of marine and terrestrial biota is considered to have continued into  
37 the initial part of the Middle Triassic (Anisian period) (Erwin, 1993; Hallam, 1991). The failure  
38 of most communities to flourish during the Early Triassic indicates the persistence of harsh  
39 conditions (Hallam, 1991; Dickins, 1993). For example, metazoan reef systems (Senowbari-

40 Daryan et al., 1993), forest ecosystems (Looy et al., 1999), and seafloor communities (Hallam  
41 and Wignall, 1997) showed recovery difficulties in the Early Triassic. Most studies investigating  
42 this period suggest that the delay in recovery may be attributed to the persistence of unfavorable  
43 climatic and/or oceanographic conditions such as an intense hothouse climate, acidification of  
44 the ocean, or ocean anoxia (Erwin, 1993; Hallam and Wignall, 1997; Wignall and Twitchewtt,  
45 2002). However, several marine faunas including conodonts and ammonoids diversified  
46 explosively in the initial part of the Early Triassic (Brayard et al., 2006, 2009; Orchard, 2007).

47 In the last decade, the focus of research has expanded from the extinction event at the  
48 Permian–Triassic (P/T) boundary to the mode and environmental conditions during ecosystem  
49 recovery in marine and terrestrial environments. The timing and key factors contributing to this  
50 recovery have been broadly studied (e.g., Knoll et al., 2007; Payne et al., 2010). Recently,  
51 Hermann et al. (2011, 2012) reported palynological and particulate organic matter data from the  
52 Early Triassic in Pakistan and South Tibet. However, few studies have discussed the  
53 environmental changes, including weathering in terrestrial environments at mid-latitudes in the  
54 Early Triassic, except for those at the P/T boundary, although definitive evidence regarding  
55 reconstruction of paleosols and faunal provinces is reported (e.g., Bourquin et al., 2011).

56 The aim of the present study is to provide constraints on environmental changes on the  
57 basis of the geochemical composition of mudstones during the Early Triassic in a well-dated  
58 succession. For this purpose, sedimentary sequences were selected from the Manang area of the  
59 Tethyan Himalayas, central Nepal (Figs. 1-A and B), which was located in the mid-  
60 paleolatitudes and faced the Neo–Tethys Ocean in the southern hemisphere (Golonka and Ford,  
61 2000). In addition, the elemental ratio of terrigenous material is very important in estimating  
62 weathering conditions in the hinterlands. Because CaO, which is a primal component of

63 carbonates, disturbs the elemental ratio of soluble–insoluble elements in terrigenous materials,  
64 mudstone geochemistry is utilized in this study.

65

## 66 **REGIONAL GEOLOGY**

67           The stratigraphy and structure of the selected sections of the Permian and Triassic  
68 sequences in the Manang area, Tethyan Himalaya, central Nepal have been studied previously  
69 (Bordet et al., 1975; Bassoullet and Colchen, 1977; Fuchs, 1977; Fuchs et al., 1988; Garzanti et  
70 al., 1994a; von Rad et al., 1994; Baud et al., 1996). The geology has been classified into the  
71 following four geological units: Permian Puchenpra Formation, “topmost biocalcarenes,”  
72 Lower Triassic Tamba Kurkur Formation, and Middle–Upper Triassic Mukut Formation.

73           These sequences are stratigraphically constrained in this area by conodont and ammonoid  
74 fossils (Bordet et al., 1975; Bassoullet and Colchen, 1977; Fuchs et al., 1988; Garzanti et al.,  
75 1994a, b; Waterhouse, 2010). The Lower Triassic Tamba Kurkur Formation in the Manang area  
76 is stratigraphically well constrained by index fossils such as conodonts (Garzanti et al., 1994a, b)  
77 and consists of the following stages (Fig. 1-C): Griesbachian, which includes the topmost  
78 biocalcarenes and first carbonate band; Dienerian, which primarily represents the first  
79 mudstone interval; Smithian, which primarily includes the second carbonate band; and Spathian,  
80 which represents the second mudstone interval and third carbonate band. Major facies and ages  
81 are described in the following paragraph.

82

### 83 ***Puchenpra Formation***

84           This formation, which was separated from the Thini Chu Group by Garzanti (1999),  
85 consists of quartz sandstone and mudstone with limestone–marl intercalations. The Kuling

86 Group of the Spiti–Zaskar area is equivalent to this formation (Garzanti et al., 1996b; Baud et  
87 al., 1996). The original definition of the Thini Chu Formation by Bodenhausen et al. (1964)  
88 included dark mudstones and white sandstones intercalated with richly bioclastic intervals of the  
89 Carboniferous age and “exceptionally ill-sorted sandstones.” Garzanti (1999) divided this  
90 formation into the Thini Chu Group of Carboniferous–lowermost Permian age and the Puchenpra  
91 Formation, which unconformably overlies the Thini Chu Group. The age of the Puchenpra  
92 Formation is considered to be from the Sakmarian/Artinskian to Wuchiapingian on the basis of  
93 calcareous benthic foraminifera (Colchen and Vachard, 1975), brachiopods, bivalves, and  
94 conodonts (Garzanti et al., 1994b). In the sample section, a 50-m-thick section below the top is  
95 cropped out.

96

#### 97 *Topmost biocalcarenites*

98 A thin, distinct unit consisting of orange-bedded biocalcarenite with dolomitized micritic  
99 groundmass appears on the Puchenpra Formation (Bassoullet and Colchen, 1977). This unit is  
100 called the “Pangjang Formation” to ascribe to the basal Triassic (Waterhouse, 1977; 1994; Hatleberg and  
101 Clark, 1984; Baud et al., 1996) or the topmost biocalcarenites as a topmost part of the Puchenpra  
102 Formation (Garzanti et al., 1994a, b). Although the age range is Kubergandian–Murgabian, as  
103 indicated by brachiopods and conodonts at the base, the mixing of macrofossils and conodonts of  
104 Permian and Triassic ages occurs at the top of this horizon (Nicora and Garzanti, 1997). Because  
105 uncertainly in age and sedimentological characters still remain, the latter informal name (Garzanti  
106 et al., 1994a) is used in the present study.

107 This unit, which is 1.1–2.5 m in thickness, consists of orange-weathered reddish  
108 biocalcareites, including brachiopods, fenestellid bryozoans, corals, crinoids, and bivalves, along  
109 with quartz grains.

110

111 ***Tamba Kurkur Formation***

112           The age of the Tamba Kurkur Formation ranges from Griesbachian to upper Spathian or  
113 Anisian, according to the biostratigraphy of ammonoids and conodonts (von Rad et al., 1994;  
114 Garzanti et al., 1994a), and conformably overlies the topmost biocalcarenes. Its thickness is 60  
115 m and is composed of marly limestone and mudstone.

116           Detailed stratigraphy has been established on the basis of conodonts, brachiopods, and  
117 ammonoids in the Manang area (Nicora et al., 1992; Garzanti et al., 1994a). The major index  
118 fossils (Garzanti et al., 1994a) and lithofacies are shown in Fig. 1-C. The lithological  
119 characteristics in the study area suggest that this Lower Triassic system can be classified into six  
120 lithological units composed of three carbonate bands (first, second, and third carbonate bands)  
121 and two mudstone intervals (first and second mudstone intervals).

122

123 ***Mukut Formation***

124           The Mukut Formation in the Manang area consists of a 200-m-thick sequence of  
125 limestone and shale from Anisian to late Carnian in age (Garzanti et al., 1994a). In the sample  
126 section, a 40-m interval in the lower part of the Mukut Formation is observed. The limestone  
127 consists of bioclastic wackestones and marlstones and includes foraminifera, ostracods,  
128 echinoderms, ammonites, and brachiopods.

129

130 ***Biostratigraphy***

131           Several important fossils have been reported (Garzanti et al., 1994a; Waterhouse, 2010)  
132 in the Lower Triassic strata near the Manang area. Waterhouse (2010) reported basal Triassic,

133 Griesbachian ammonoids *Otoceras woodwardi* and *Ophiceras tibeticum* from the topmost  
134 biocalcarenes. The first carbonate band of the Tamba Kurkur Formation (Fig. 1-C-1) yielded  
135 Griesbachian conodonts, including *Hindeodus* cf. *typicalis* Sweet, 1970, and *Gondolella carinata*  
136 Clark, 1959. In the first mudstone interval (Fig. 1-C-2), early Dienerian conodonts, *Gondolella*  
137 *nepalensis* Kozur & Mostler, 1976, were recovered from the lower part and Late Dienerian–early  
138 Smithian *Nepsothodus pakistanensis* Sweet, 1970 were found in the upper part in the Manang  
139 area. The second carbonate band (Fig. 1-C-3 and 4) yielded *N. pakistanensis*, *N. waageni* Sweet,  
140 1970, and *Gondolella sweeti* Bender, 1970, suggesting a Smithian age. The thin mudstone  
141 interval (Fig. 1-C-5) between the second and third carbonate bands yielded *Gondolella* aff.  
142 *Jubata*, Bender, 1970, suggesting a late Smithian age. From the third carbonate band (Fig. 1-C-6),  
143 late Smithian–late Spathian conodonts, *G.* aff. *Jubata*, *Neospathodus homeri* Bender, 1970,  
144 *Gladigondolella carinata* Bender, 1970, and *Neospathodus spathi* Bender, 1970, have been  
145 observed. *Gondolella timorensis* Bender, 1970, which was found at the top of the third carbonate  
146 band (Fig. 1-C-7) by Garzanti et al. (1994a), indicates the late Spathian (Goudemand et al., 2012).  
147 The lower part of the Mukut Formation in the Manang area yielded Early Anisian fossils  
148 (Garzanti et al., 1994a) such as conodonts: *Gondolella regalis* (Mosher, 1970); brachiopods:  
149 *Punctospirella stracheyi* and “*Dielasma*” *himalayanum*; and ammonoids: *Hoolandites* sp.

150         Recently, the Tulong Formation in South Tibet, which has very similar lithology as the  
151 Tamba Kurkur Formation in the Manang area, was revised by Brühwiler et al. (2009) using  
152 ammonoids and conodonts. The stratigraphy of the Tulong Formation suggests that the first  
153 carbonate ranges from middle to late Smithian and the second mudstone interval is Spathian  
154 (Brühwiler et al., 2009). Thus, the chronological framework of the Tamba Kurkur Formation in

155 the Manang area proposed by Garzanti et al. (1994a) has been changed to include the above  
156 recent revision of the Triassic biostratigraphy.

157

## 158 **FACIES DESCRIPTION IN THE SAMPLE SECTION**

159 On the northern cliff of Manang village, the Lower Triassic deposits are well exposed  
160 and are easily accessible. The continuous sequence from the Permian Puchenpra Formation to  
161 Upper Triassic Mukut Formation was reported by Garzanti *et al.* (1994a), although the Mukut  
162 Formation, which is distributed over an altitude of 4500 m on the cliff, is not accessible. The  
163 abovementioned strata have an east–west trending strike and are nearly horizontal or gently  
164 dipping to the north–northwest (Fig. 2-A).

### 165 ***Permian Unit***

#### 166 **Puchenpra Formation**

167 The Puchenpra Formation in this section consists of approximately 50 m of alternating  
168 beds of sandstone and mudstone. The sandstone bed shows troughs and hummocky cross-  
169 stratification (HCS) (Fig. 2-B). The mudstone is intensely bioturbated and includes large U- and  
170 I-shaped burrows.

### 171 ***Triassic Unit***

172 The Triassic unit consists of topmost biocalcarenites and the Tamba Kurkur and Mukut  
173 Formations. The Triassic Tamba Kurkur Formation is divided into five subunits (Fig. 1-C). The  
174 lower two units are mostly similar to that reported by von Rad et al. (1994) and Baud et al.  
175 (1996) for the Jomsom area, although the upper three units differ with respect to thickness and  
176 lithofacies.

#### 177 ***Topmost biocalcarenites***

178 Orange dolomitic sandy limestone, which is 1.1 m in thickness, lies above the  
179 uppermost part of the Puchenpra Formation (TB in Figs. 2-A and C). This dolomitic sandy  
180 limestone consists of stratified bioclastic wackestones and includes fine sand-sized quartz grains  
181 and calcareous bioclasts such as brachiopods, bryozoans, ostracods, small foraminiferans, and  
182 crinoids (Fig. 2-D). Individual beds show massive structures, including oversize clasts of  
183 bryozoans and brachiopods. The contact between the base of this unit and the mudstone of the  
184 uppermost part of the Puchenpra Formation shows a slightly deformed undulating plane,  
185 probably suggesting loading or erosional structures (Fig. 2-C).

186

## 187 **Tamba Kurkur Formation**

### 188 *First carbonate band*

189 The first carbonate band, which is located at the base of the Tamba Kurkur Formation,  
190 consists of orange- and gray-colored dolomitic limestone 3 m in thickness. At the base of this  
191 interval, several layers of 30–50-cm-thick orange dolomitic limestone rest on the orange  
192 dolomitic sandy limestone of the topmost biocalcarenes (TB in Fig. 2-C). Although the  
193 boundary between the topmost biocalcarenes and Tamba Kurkur Formation is a planar contact,  
194 there is no clear evidence of erosion. The limestone of the Tamba Kurkur Formation consists of  
195 fine bioclastic wackestone and mudstone, including thin-shelled bivalves and small ammonites  
196 without quartz grains. Although the gray wackestone that rests on the abovementioned bioclastic  
197 wackestone and mudstone is recrystallized and dolomitized, it still includes thin-shelled bivalves,  
198 small ammonoids, and radiolarians (Fig. 2-H). These limestone layers, along with the uppermost  
199 Permian (Tc1 + TB in Fig. 2-A), show distinctive marker horizons, which make them easily  
200 recognizable in the outcrop.

201            *First mudstone interval*

202            The first mudstone interval consists of 30-m-thick dark red–black mudstones. These  
203 mudstones contain thinly bedded, alternating layers of 1–10-mm-thick red–purple mudstone and  
204 black mudstone (Fig. 2-E). Small burrows approximately 5 mm in diameter are rarely observed  
205 (Fig. 2-F). The mudstone layer shows grading (Fig. 2-G). In the lower part of this mudstone  
206 interval, several thin-bedded orange dolomitic limestone layers are intercalated.

207            *Second carbonate band*

208            An 8–10 m thick, second carbonate band consists of gray–yellow nodular limestone  
209 with intercalated thin gray mudstone layers. In the lower part of this carbonate band, the yellow  
210 limestone is composed of wackestone–packstone containing numerous fragments of thin-shelled  
211 bivalves. In the upper part, the gray–yellow nodular limestone consists of wackestone–packstone  
212 with numerous ammonoids and thin-shelled bivalves (Fig. 2-I). The uppermost horizon of this  
213 unit consists of a peculiar packstone–grainstone containing large numbers of ostracods with  
214 ammonoids, gastropods, sponge spicules, and crinoids.

215            *Second mudstone interval*

216            The mudstone temporally intercalated between the second and third carbonate bands has  
217 no characteristic sedimentary structures such as wave ripple or HCS. This interval is a 20–30-  
218 cm-thick dark gray mudstone between the second and third carbonate bands. The thin beds of  
219 calcareous mudstone or muddy limestone are rarely intercalated. This mudstone consists of dark  
220 gray shales with weak bioturbation. Examination under a microscope reveals that the original  
221 thin lamination is preserved; however, such lamination is frequently disturbed by small burrows.

222            *Third carbonate band*

223           The third carbonate band consists of a 12-m-thick gray nodular limestone that includes  
224 ammonite-bearing gray nodular limestone and alternating beds of gray limestone and mudstone  
225 in the lower part. The gray limestone in the lower part is a wackestone containing ostracods and  
226 thin-shelled bivalves with some ammonoids (Fig. 2-J). The upper part consists of nodular gray  
227 limestone and reddish gray nodular limestone intercalated with thin gray and red calcareous  
228 mudstone. In this part, the limestone clearly includes small ammonites, ostracods, radiolarians,  
229 and thin-shelled bivalves along with fine sand-sized quartz grains. However, as previously  
230 mentioned, this horizon is intensely deformed and recrystallized (Fig. 2-K).

231           In the uppermost part of this unit, stratified limestone with a few mudstone  
232 intercalations gradually grade into thin-bedded alternating layers of limestone and mudstone in  
233 the lower part of the Mukut Formation (Fig. 2-L).

234

### 235 *Interpretation of sedimentary environments*

236           A total of five units reflect the main depositional environments through sedimentary  
237 structure, grain size, and bioturbation amount.

### 238 *Puchenpra Formation*

239           The sedimentary facies of the Puchenpra Formation indicate offshore environments  
240 located between storm and fair-weather wave bases. A large amount of bioturbation is  
241 appropriate for this environment.

### 242 *Topmost biocalcarenites*

243           In this unit, the lack of sedimentary structures that present in the shallow marine  
244 environment suggests deposition below storm wave base. Because this unit consists of an ill-

245 sorted mixture of quartz grains and small and large fragments of bioclasts, they may be lag  
246 deposits with rapid transgression.

247 ***First carbonate band***

248 This unit consists of massive wackestone and lime–mudstone with an absence of shallow  
249 marine sedimentary structures. This facies indicates a lower-energy environment offshore or  
250 deposition below the storm wave base.

251 ***First mudstone interval***

252 Bioturbation is rare and small-scale laminations have been completely preserved in this  
253 unit. Grading and absence of shallow marine sedimentary structures indicate deposition in an  
254 offshore environment or that below the storm wave base.

255 ***Second carbonate band***

256 The lithofacies of this unit is characterized by fossiliferous wackestone. In the lower part,  
257 a concentration of thin-shelled bivalve clasts reflects an episode of gregarious deposition at the  
258 sea bottom. In the upper part, nektonic and benthic fossils such as ammonoids and bivalves are  
259 prevalent as constituents of the limestone, indicating deposition in hemiplegic environments.

260 ***Second mudstone interval***

261 Minor carbonate intercalations suggest a temporary decrease in carbonate productivity or  
262 of the input of prevailing terrigenous material. The absence of shallow marine sedimentary  
263 structures suggests deposition below the storm wave base or within a very short time interval.

264 ***Third carbonate band***

265 The prevailing wackestone with benthic fossils and a lack of typical shallow marine  
266 sedimentary structures implies deposition in low-energy hydraulic bottom environment. The  
267 remaining structure of bivalve concentration suggests deposition in a pelagic environment. The

268 intercalated red calcareous mudstone in the upper part may be correlated to the Ammonitico  
269 Rosso facies observed in the Jomsom area, central Nepal (von Rad *et al.*, 1994) and Tulong area,  
270 South Tibet (Brühwiler *et al.*, 2009). The sandy wackestone containing quartz grains may be  
271 indicative of temporal change of transporting patterns of sediments and rapid predominance of  
272 terrigenous material.

273

#### 274 **DESCRIPTION OF MUDSTONES IN THE SAMPLE SECTION**

275 The mudstones of the Tamba Kurkur Formation can be described as weakly  
276 metamorphosed gray–black mudstones with lithologies ranging from claystone to coarse  
277 siltstone. Small-sized burrows ranging from 2–5 mm in diameter are occasionally included.

278 The quartzose and feldspathic grains in the mudstones are predominantly angular to  
279 subangular. Graded bedding is mostly obscured due to recrystallization and weak deformation.  
280 Deformation by compaction is observed, as evidenced by commonly deformed calcareous  
281 biogenic spherical grains.

282 The characteristics of these mudstones, in conjunction with petrographic observations  
283 revealed under a microscope, indicate the following characteristics:

284 1. The mudstones in the Puchenpra Formation are matrix-supported siltstones with very  
285 fine quartz grains. Quartz grains are angular–subangular. In addition, biotite, white mica, and  
286 tourmaline grains are included. Calcareous grains originating from the fragments of calcareous  
287 organisms are frequently included. Authigenic rhombohedral dolomite grains are commonly  
288 observed.

289 2. The mudstone in the first mudstone interval of the Tamba Kurkur Formation is fine  
290 siltstone to claystone with quartz grains 0.005–0.35 mm in diameter, with an average grain size

291 of 0.017–0.029 mm (Fig. 3-a and b). Although slaty cleavage is present, the sample contains  
292 detrital muscovite flakes that are locally oblique to the cleavage of the metamorphic mica  
293 minerals. Occasionally, detrital calcite or biogenic calcite grains are included. Authigenic  
294 dolomite and pyrite grains can be observed. Several mudstones, which consist of coarse-grained  
295 minerals of metamorphic chlorite and mica, are intercalated in the first mudstone interval.

296 3. The mudstones in the second carbonate band are composed of thin layers intercalated  
297 with limestone. These mudstones are fine siltstones to claystones containing fine quartz and  
298 feldspar grains of 0.005–0.04 mm in diameter, with an average grain size of 0.014–0.021 mm  
299 (Fig. 3-c and d). Detrital muscovite is rare but present. The abundance of authigenic dolomite  
300 grains increases.

301 4. The mudstones derived from the third carbonate band are fine siltstones to claystones  
302 with quartz grains of 0.005–0.03 mm in diameter. Detrital muscovite is rare. Authigenic  
303 dolomite grains are included, along with metamorphic chlorite and mica minerals.

304

## 305 **RESULTS OF GEOCHEMISTRY**

### 306 *Carbon isotope record in the Lower-Middle Triassic carbonate in Manang area*

307 In this study, Early Triassic  $\delta^{13}\text{C}_{\text{carb}}$  data from the Manang area, which is constrained by  
308 conodont stratigraphy (Garzanti *et al.*, 1994a), are presented. Because samples of the Lower  
309 Triassic Tamba Kurkur Formation display very low  $\delta^{18}\text{O}_{\text{carb}}$  values, with an average value of  
310  $-12\text{‰}$ , and unusually low  $\delta^{13}\text{C}_{\text{carb}}$  values during the Griesbachian to late Smithian, with an  
311 average value of  $-1.5\text{‰}$  (Fig. 3), one may question whether the primary marine characteristics  
312 have been preserved or whether the measured isotope values represent diagenetic effects.  
313 Garzanti *et al.* (1994a, 1998) suggests that there may be some significant influence from

314 diagenesis and metamorphism in the Permian/Triassic sections in South Tibet and the Manang  
315 area. In particular, significant metamorphic conditions are indicated by illite crystallinity,  
316 vitrinaite reflectance, and conodont color in the Manang area (Garzati *et al.*, 1994a). A  
317 comparison between the  $\delta^{13}\text{C}_{\text{carb}}$  and  $\delta^{18}\text{O}_{\text{carb}}$  profiles show that  $\delta^{18}\text{O}_{\text{carb}}$  varies with  $\delta^{13}\text{C}_{\text{carb}}$   
318 values in the second carbonate band, which suggests some influence of diagenesis on the isotope  
319 records. It is widely known that oxygen isotope values are more easily altered during diagenesis  
320 than carbon isotope values (Marshall, 1992). In the Tethyan Himalayas, very low  $\delta^{18}\text{O}_{\text{carb}}$  values  
321 have been found and explained by thermal metamorphism of up to 300 °C (Baud *et al.*, 1996;  
322 Atudorei and Baud, 1997). However, it is known that relative changes in the carbon isotope  
323 excursion are preserved in low-grade metamorphic conditions and that the signal can be  
324 correlated with other records for the same time span, as shown in other Himalayan sections  
325 (Baud *et al.*, 1996; Atudorei and Baud, 1997; Brühwiler *et al.*, 2009), although oxygen isotope  
326 values have been altered during diagenesis and metamorphism.

327 A bulk carbonate carbon isotope curve generated for the lower and upper parts of the  
328 Tamba Kurkur Formation shows a positive excursion. This correlation is discussed subsequently.

### 329 ***Inorganic geochemistry***

330 The samples studied have been categorized into the following six groups: (1) mudstones  
331 intercalated in the Permian Puchenpra Formation, (2) mudstones in the topmost biocalcarenites,  
332 (3) the first mudstone interval, (4) mudstones intercalated in the second carbonate band, (5)  
333 mudstones in the third carbonate interval of the Tamba Kurkur Formation, and (6) mudstones in  
334 the lowermost horizon of the Mukut Formation.

335 In Fig. 5,  $\text{Al}_2\text{O}_3$ ,  $\text{Fe}_2\text{O}_3^*$ ,  $\text{MnO}$ , and  $\text{MgO}$  concentrations generally show strong negative  
336 correlations with  $\text{SiO}_2$ . The wide compositional variety of the mudstones in the first mudstone

337 interval tends to mask the characteristics of the mudstones in the other stratigraphic units. The  
338 mudstones in Permian and topmost biocalcarenes show lower concentrations of  $\text{Al}_2\text{O}_3$ ,  $\text{MgO}$ ,  
339  $\text{Fe}_2\text{O}_3^*$ , and  $\text{P}_2\text{O}_5$  and higher concentrations of  $\text{Na}_2\text{O}$  and  $\text{CaO}$ , which are caused by its sandy  
340 nature with extended feldspar grains. The mudstones from the first mudstone interval appear to  
341 have a similar elemental composition to the third carbonate band and Mukut Formation. This  
342 interval also includes two samples that are characterized by high  $\text{Fe}_2\text{O}_3^*$  and  $\text{MnO}$   
343 concentrations and low  $\text{SiO}_2$  and  $\text{Na}_2\text{O}$  concentrations. This is due to the high content of mica  
344 and Fe-oxide minerals. The mudstones in the second carbonate band show higher  $\text{SiO}_2$ ,  $\text{CaO}$ ,  
345  $\text{Na}_2\text{O}$ , and  $\text{P}_2\text{O}_5$  concentrations and lower concentrations of  $\text{Al}_2\text{O}_3$ ,  $\text{Fe}_2\text{O}_3^*$ , and  $\text{MgO}$  compared  
346 with other mudstones, which distinguishes their siliceous nature with lower alumina and iron  
347 contents.

#### 348 ***Rare Earth Elements (REEs)***

349 The average REE pattern shows that all six groups of sediments have similar chondrite-  
350 normalized REE patterns with overlapping abundances (Fig. 6). The concentrations show an  
351 increase in order of magnitude from Permian sediments and the second carbonate band to the  
352 first mudstone interval, the third carbonate band of the Triassic Tamba Kurkur Formation, and  
353 the Mukut Formation within a very narrow range. All groups of sediments show fractionated,  
354 parallel to subparallel REE patterns, with  $(\text{La}/\text{Yb})\text{N}$  ratios ranging from 13.8 to 16.9  
355 (Supplemental Table 1). Both LREE and HREE show variable fractionation with  $(\text{La}/\text{Sm})\text{N}$   
356 values ranging from 4.8 to 5.1. The  $(\text{Gd}/\text{Yb})\text{N}$  value ranges from 1.6 to 1.9. This indicates that  
357 the REE content among various groups of mudstones varies within a narrow limit and that their  
358 fractionation is not large because the differences in both  $(\text{La}/\text{Sm})\text{N}$  and  $(\text{Gd}/\text{Yb})\text{N}$  ratios are of  
359 the same order of magnitude. All samples show consistently uniform negative Eu anomalies. The

360 samples of the Tamba Kurkur and Mukut Formations do not show Ce anomalies, although  
361 Permian and topmost biocalcarenites samples show weakly positive Ce anomalies. Note that the  
362 REE patterns of the Triassic mudstones are similar to those of post-Archean Australian average  
363 shale (PAAS; Fig. 6), including the negative Eu anomalies. The patterns, which are comparable  
364 to PAAS, imply that the detrital material was derived from a common continental source.

### 365 *Chemical weathering degree*

366 It is widely accepted that marine mudstones preserve information about their sedimentary  
367 environments and hinterlands (Potter et al., 2005). In particular, information regarding terrestrial  
368 environments such as climatic conditions can be estimated by examining the paleoweathering of  
369 these rocks to determine the extent of chemical weathering. In general, warm and wet climates  
370 provide ample moisture for the hydrolysis of common elements such as Ca, Mg, Na, and K from  
371 surface rocks and soil minerals in the most weathered parts of the soil. On the contrary, arid and  
372 cool climates contain insufficient water for flushing cations from the soil into the groundwater.  
373 As a result, alkaline earth cations such as Ca and Mg remain within the soil as subsurface  
374 nodules (Retallack, 2005). Thus, several geochemical indices for estimating weathering effects,  
375 including CIA (Nesbitt and Young, 1982), have been used to derive transfer functions for  
376 paleoclimate estimation from sediments. In this study, CIA values along with  $\text{Al}_2\text{O}_3/\text{Na}_2\text{O}$  ratios  
377 are used to estimate degrees of weathering.

378 In the Manang area, which includes most of the Upper Permian to Lower Triassic  
379 sequence, the geochemical indices suggest significant vertical compositional variations (Fig. 7  
380 and Supplemental Table 2). Most Lower Triassic sediments in the first mudstone interval of the  
381 Tamba Kurkur Formation show fairly constant CIA values and  $\text{Al}_2\text{O}_3/\text{Na}_2\text{O}$  ratios of 72–78  
382 (average = 76.2, s.d. = 1.9) and 21–31 (average = 26.8, s.d. = 2.7), respectively. Four samples

383 obtained from the upper part of the first mudstone interval and from the lower part of the second  
384 carbonate band show the CIA values of 67–69 (average = 68.1, s.d. = 1.3) and Al<sub>2</sub>O<sub>3</sub>/Na<sub>2</sub>O ratios  
385 of 12–16 (average = 13.6, s.d. = 1.9; Fig. 7). From the third carbonate band of the Tamba Kurkur  
386 Formation to the Mukut Formation, the CIA values and Al<sub>2</sub>O<sub>3</sub>/Na<sub>2</sub>O ratios were relatively higher,  
387 reaching 80 and 30, respectively. The index of compositional variability (ICV) values, which  
388 indicate the maturity of fine alumino–siliciclastic materials (Cox et al., 1995), were relatively  
389 constant with some minor fluctuations throughout the sample section, including exceptional two  
390 values in the lowest horizon of the section (<0.5) and in the first mudstone interval (>1.5).

### 391 *Grain-size analysis*

392 In the Tamba Kurkur Formation, the average grain size measured from mudstones ranged  
393 from 0.017 mm in the lower part of the first mudstone interval to 0.029 mm in the middle part of  
394 the same interval. However, a sample from the uppermost part of the Puchenpra Formation  
395 contained a coarser grain size of 0.043 mm. In contrast, samples from the upper part of the first  
396 mudstone interval and third carbonate band showed no significant differences in grain size.

397

## 398 **DISCUSSION**

### 399 *Carbon isotope chemostratigraphy*

400 The Triassic bulk carbon isotope data of the study area correlates to the Lower Triassic  
401 Tamba Kurkur Formation in the Tulong area, South Tibet (Brühwiler et al., 2009). These strata,  
402 which are distributed in the Tulong area, had been emended from the Tulong Formation by  
403 Garzanti *et al.* (1998). The lithostratigraphy of the Tamba Kurkur Formation in the Manang area  
404 is closely correlated to that of the Tulong area. The generated excursion of the Tamba Kurkur  
405 Formation in the Manang area shows considerable correlation with that in the Tulong area. On

406 the basis of the correlation potential of the isotope curve between both areas, we differentiated  
407 four excursion types: prominent negative excursion ( $-5\text{‰}$ ) around the P/T boundary, significant  
408 positive excursion ( $+3\text{‰}$ ) at the Smithian–Spathian boundary (horizon A in Fig. 1-C), positive to  
409 negative excursion around the lower Spathian (horizon B in Fig. 1-C), and gentle positive  
410 excursion in the upper Spathian (horizon C in Fig. 1-C). Thus, combined correlations based on  
411 conodont presence and carbon isotope records in the Tulong area suggest that the approximate  
412 age of the horizon with low CIA values and  $\text{Al}_2\text{O}_3/\text{Na}_2\text{O}$  ratios is upper Smithian.

413

#### 414 ***Record of sea level change***

415 The sediments of the Lower Triassic show very thin thicknesses of approximately 50 m  
416 from Griesbachian to Spathian, representing a highly condensed sequence. The Lower Triassic  
417 sequence in the Tethys Himalayas is characterized by similar condensed sequence (von Rad et al.,  
418 1994; Baud et al., 1996; Galfetti et al., 2007a; Brühwiler et al., 2009). Although the thickness of  
419 stratigraphic succession always demands the consideration of tectonism and eustasy, the  
420 condensed sequence in a wide area of the Tethys Himalayas indicates that the sedimentation  
421 represents global sea level change that occurred during a transgressive stage. The cyclic  
422 sedimentation of carbonates and mudstone can be considered as a reflection of eustatic sea level  
423 changes in the sample section. The carbonate bands in the sample section could be correlated to  
424 the transgressive phase of the global cycle chart by Haq et al. (1988), Hardenbol et al. (1998),  
425 and Haq and Shutter (2008) (Fig. 1-C). The topmost biocalcarenes includes a large amount of  
426 bioclasts such as corals, bryozoans, and brachiopods with quartz grains, which is indicative of  
427 sediment reworking. Moreover, the transition from the Permian Puchenpura Formation to the  
428 topmost biocalcarenes is marked by a sharp erosional surface and a change of lithofacies from

429 alternation of sandstone and mudstone to dolomitic orange sandy carbonate. Thus, the onset of  
430 the sedimentation of the Triassic system is interpreted to be related to starvation and subaqueous  
431 reworking of the sediments during early Griesbachian rapid transgression. The first carbonate  
432 band of the Tamba Kurkur Formation, directly overlying the topmost biocalcarenes, could be  
433 correlated to a highstand system tract in Griesbachian highstand. The first mudstone interval,  
434 which is rich in terrigenous material, likely suggests the lowstand systems tract in Dienerian and  
435 early Smithian because the input of terrigenous material was increased due to exposure of the  
436 shelf. The second and third carbonate bands could be correlatable to the late Smithian highstand  
437 and middle–late Spathian transgressive highstand system tracts, respectively. The second  
438 mudstone interval is likely correlative to the lowstand system tract with early Spathian regression.

439         Similar eustatic cycles are known from the Canadian Sverdrup Basin (Embry, 1997) and  
440 are recognized in many other sections such as the Western Tethys (Gianolla and Jacquin, 1998),  
441 Barents Sea (Gianolla and Jacquin, 1998; Skjold et al., 1998), and the Salt and Sugar Ranges  
442 (Hermann et al., 2011).

443

#### 444 ***Provenance***

445         Some elements, and particularly their ratios, are useful indicators of provenance because  
446 they are least affected by processes such as weathering, transport, and sorting. In particular, the  
447 common immobile elements such as REEs have been found to be useful (Taylor and McLennan,  
448 1985). In addition, REEs are believed to preserve the composition of the source rock in which  
449 they are carried by the detrital component (Taylor and McLennan, 1985; Wronkiewicz and  
450 Condie, 1987). These elements indicate specific conditions in igneous petrogenesis. Hence,  
451 normalized REE patterns suggest low (La/Yb)<sub>N</sub> ratios and little or no Europium (Eu) anomalies

452 in mafic rocks. In contrast, felsic rocks have higher (La/Yb)<sub>N</sub> ratios and significant Eu anomalies  
453 (Taylor and McLennan, 1985). The chondrite normalized REE patterns in all sediments in each  
454 unit of the sample section have been exactly fractionated and show significant Eu anomalies that  
455 suggest felsic source rock provenance.

456         The chondrite normalized REE patterns of the mudstones have been compared with those  
457 of PAAS. All patterns from the Triassic stratigraphic unit shown in this study are similar to that  
458 of the average shale. Moreover, these mudstones show enrichment of REE in comparison with  
459 PAAS. This feature implies that the REE patterns in each Triassic unit are neither affected by  
460 selective sorting in sedimentary process nor derived from different source rock types. Therefore,  
461 the enrichment requires a concentration of REE without significant changes in REE patterns.  
462 This process may be related to the elemental concentration due to chemical weathering, which  
463 leads to chemical leaching of soluble elements, leaving the insoluble elements in the hinterland.  
464 Cullers et al. (1987) indicated that the enrichment of REE in the clay fractions of stream  
465 sediments is relative to source rocks. They concluded that the REEs are released from primary  
466 minerals and taken up by the clay size fraction during chemical weathering of the source rock.

467

#### 468 *Paleoweathering record*

469         The sediment chemistry indicates the results of the influence of transport processes,  
470 source rocks, and weathering (McLennan et al., 1983). The marine sediments reflect the  
471 characteristics of weathered source rocks and effects related to fluvial transportation and  
472 depositional environments. Thus, the compositions of initial detritus have been affected and  
473 modified by multiple factors such as climate, selective sorting in fluvial and marine processes,  
474 diagenesis, and metamorphism. Of these factors, the effects of diagenesis and metamorphism,

475 which could not be completely excluded, are considered to be insignificant in this study because  
476 diagenetic and metamorphic alternations at specific horizons were not found in the sample  
477 section, which is indicated by the specific chemical composition and occurrence of authigenic  
478 mineral assemblages observed under a microscope. Although authigenic dolomite, metamorphic  
479 mica minerals, and iron minerals are observed in these rocks, biased bulk rock chemistry due to  
480 such specific mineral formation is not found.

481 The evolution of chemical weathering of sediments can reflect changes in the  
482 environment and may record the history of environmental changes in the depositional basin. It is  
483 widely accepted that weathering profiles have a significant effect on the geochemical  
484 compositions of siliciclastic sediments (Nesbitt and Young, 1982). Accordingly, the composition  
485 and subsequent CIA values of sediments can indicate the intensity of chemical weathering  
486 associated with climatic conditions (Nesbitt and Young, 1982; Fedo *et al.*, 1997).

487 The uppermost horizon of the Permian strata of the Puchenpra Formation exhibited  
488 relatively variable CIA values of 67–73 and  $\text{Al}_2\text{O}_3/\text{Na}_2\text{O}$  ratios from 11 to 19. The topmost  
489 biocalcarenes represented CIA values of 69 and  $\text{Al}_2\text{O}_3/\text{Na}_2\text{O}$  ratios of approximately 16.  
490 However, most samples from the Tamba Kurkur Formation showed CIA values of 72–79, which  
491 are similar or slightly higher than those of PAAS and North American shale composite (NASC),  
492 i.e., 70–75 for average shales (Taylor and McLennan, 1985). These results suggest intense  
493 weathering conditions during the Griesbachian to Dienerian and early Smithian, even in the  
494 middle paleolatitudes, due to a hot–warm, humid climate, as indicated by previous  
495 paleontological and geochemical studies (Worsley *et al.*, 1994; Kidder and Worsley, 2004). In  
496 contrast, the samples from the upper part of the first mudstone interval to second carbonate band  
497 (late Smithian) show low CIA values of 67–69, which suggests deposition in environments in

498 which intense weathering was absent in the hinterland. The CIA values show an obvious degree  
499 of hydrolysis of alkaline cations because the vertical change in these values is correlated with  
500  $\text{Al}_2\text{O}_3/\text{Na}_2\text{O}$  ratios. The ICV values show no correlation with the CIA values, which implies that  
501 the effect of selective removal and addition barely contributes to the change in these values.  
502 However, the CIA values and  $\text{Al}_2\text{O}_3/\text{Na}_2\text{O}$  ratios may be affected by the sediment grain size. The  
503 grain-size distribution in the selected samples shows no relationship to the CIA values and  
504  $\text{Al}_2\text{O}_3/\text{Na}_2\text{O}$  ratios, with the exception of one sample from the topmost biocalcarenites. In  
505 addition, the average grain size of sample no. 83, which has a low CIA value, is similar to that of  
506 sample no. 86, which has a high CIA value. Thus, the CIA values are considered to reflect the  
507 degree of hydrolysis in the source area because the grain size change among these samples is not  
508 linked with either the CIA values or  $\text{Al}_2\text{O}_3/\text{Na}_2\text{O}$  ratios.

509 The A–CN–K ternary diagram indicates the weathering trend line (Fig. 8). The groups of  
510 the Triassic Tamba Kurkur Formation with Permian mudstones all plot along the same  
511 weathering trend line, which implies that they are derived from a common lithology.

512

### 513 *Factors affecting chemical composition of sediments*

514 The dominant controlling factors that affect the degree of paleoweathering are considered  
515 to be tectonic movements and climatic conditions in the source and depositional regions on land.  
516 Active tectonic forces, such as uplift of the hinterland, would sufficiently suppress chemical  
517 weathering by enhancing physical weathering and rapid accumulation. However, the lack of  
518 significant change in grain size and lithofacies indicates that there was no significant ongoing  
519 tectonic movement such as rapid uplift of the hinterland. Furthermore, a temporal decrease in the  
520 degree of paleoweathering is present from the upper part of the first mudstone interval to the

521 second carbonate band with a return to the previous degree of paleoweathering in the third  
522 carbonate band.

523         Moreover, it is well known that relative sea level change affects the degree of chemical  
524 weathering. In particular, topographically flat regions on land, such as coastal plains and large  
525 flood plains, are strongly affected by relative sea level change (Cotton, 1963; Wright, 1970). In  
526 addition, the composition of sediments along the continental margin is controlled by relative sea  
527 level change with local modifications of topography occurring as a result of changes in the  
528 catchment area and sediment supply system (Curzi et al., 2006). Moreover, the change in the  
529 degree of mixing and reworking of initial sediments as a result of either a reduction or expansion  
530 in accommodation space due to topographical changes induced by sea level changes is a  
531 considerable factor to cause a compositional change of sediments (Strand, 1998; Muller, 2001).

532         In this study, the horizon showing a low degree of paleoweathering (shaded horizon in  
533 Fig. 7) is located at the transition into the first mudstone interval and second carbonate band. In  
534 this case, there is no relationship between lithology and the degree of paleoweathering. When  
535 carbonate deposition could be correlated with the transgressive or highstand stage of sea level  
536 change, it is considered that the horizon showing low paleoweathering degree has no connection  
537 to sea level change.

538

### 539 ***Implication of the horizon showing low paleoweathering degree***

540         Climate conditions in the source area are important factors to control the degree of  
541 chemical weathering (Nesbitt and Young, 1982; Taylor and McLennan, 1985). In general,  
542 intensely weathered sediment tends to be provided from areas in a period in which wet, humid,  
543 and warm climates prevailed; in contrast, compositionally less mature sediment tends to be

544 supplied by arid and cool periods (Prins and Postma, 2000; Tebbens and Veldkamp, 2000;  
545 Muller, 2001; Yang et al., 2002, 2008; Boulay *et al.*, 2003; Curzi et al., 2006). Extended  
546 chemical weathering of flood plain deposits in humid subtropical climates can modify the initial  
547 sediment composition through reactions with groundwater (Singh, 2009). Even in mountainous  
548 regions, topographic changes leading to changes in sediment composition can be related to  
549 climate change (Yang et al., 2002).

550         Therefore, in the presented Manang section, the characteristic horizon showing a  
551 decrease in chemical weathering without substantial changes in the source rocks is likely to be  
552 closely linked to an abrupt change of weathering conditions such as climatic changes in the late  
553 Smithian. The Early Triassic climate was determined on the basis of paleoclimatic indicators,  
554 which generally suggest that in the Griesbachian interval, an extremely hot and uniform climate  
555 prevailed, whereas the Dienerian–Smithian period appears to have had steady and warm climatic  
556 conditions (Kidder and Worsley, 2004). Hermann et al. (2011, 2012) indicated a climate change  
557 from humid to dryer conditions around the Smithian–Spathian boundary in the Salt and Surghar  
558 Ranges in Pakistan and Tulong in Tibet. Moreover, Romano et al. (2013) reconstructed  
559 paleotemperature records on the basis of oxygen isotopes from fossil conodonts in the Salt Range,  
560 and they concluded that the middle–late Smithian was under extremely warm conditions. The  
561 chemical weathering data obtained here shows an abrupt decrease in the degree of weathering  
562 during the late Smithian. Conditions were more arid, indicating that this environmental change  
563 did not persist. These results imply that the climatic environment in the Early Triassic period on  
564 the northern margin of Gondwana was unstable.

565         Indications of climate change around the Smithian–Spathian period have been observed  
566 in several areas including the western Neo–Tethys (Bourquin *et al.*, 2009) and Spitsbergen in

577 Norway (Galfetti *et al.*, 2007b). An abrupt increase in immature siliciclastic materials during the  
578 Smithian period has been clearly recorded in a platform carbonate sequence of the western  
579 Paleo–Tethys (Stefani *et al.*, 2010). Because climatic patterns are extremely complicated and  
570 vary in detail, several mechanisms could play a role in the local climate; therefore, there may be  
571 a disconnection between global and local climate regimes. In addition, the sampling interval was  
572 uneven in the second and third carbonate bands, and the duration of the dryer period remains  
573 undetermined. Thus, caution should still be maintained while extrapolating our results for the  
574 Triassic sequence in the Tethyan Himalayas to a global scale. The dispersed records of the  
575 Triassic climate should be combined near the northern Gondwana to Tethyan margins, though it  
576 is apparent that significant climate changes occurred in the late Smithian.

577

## 578 **CONCLUSIONS**

579 This study investigated the geochemistry of the Early Triassic mudstones in the Tethyan  
580 Himalayas in central Nepal. On the basis of lithofacies, the sediments are considered to have  
581 been deposited in offshore to hemiplegic environments. The chondrite normalized REE diagram  
582 for the Permian to Triassic mudstones suggests that the uniformity can be correlated to PAAS  
583 and a derivation from continental provenance. The paleoweathering index of the Early Triassic  
584 Griesbachian–early Smithian mudstones indicates a sediment supply from strongly weathered  
585 sources with CIA values of 76–81 (Nesbitt and Young, 1982). However, the mudstones in the  
586 late Smithian show weakly weathered sources with the CIA values of 68–74. The change in the  
587 degree of chemical weathering in this Lower Triassic sequence shows no dependence on  
588 lithology, grain size, or ICV values (Cox *et al.*, 1995). This significant change in the degree of  
589 paleoweathering suggests an obvious change in the climatic conditions that caused continental

590 weathering in the late Smithian. Although climatic patterns are extremely complicated, the  
591 observed abrupt decrease in continental weathering suggests a dryer condition and an unstable  
592 climatic environment in the Early Triassic period on the northern margin of Gondwana.

593

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605

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## 822 **FIGURE CAPTIONS**

823 Fig. 1 A: Simplified geological map of the Himalayan area; B: Index map of the sample section  
824 in the Manang area, Tethyan Himalayas; C:  $\delta^{13}\text{C}_{\text{carb}}$  record from the Tamba Kurkur  
825 Formation in the Manang area and the correlation with the Tulong area in South Tibet  
826 (Brühwiler et al., 2009) and the Losar area, Spiti, India (Galfetti et al., 2007a). Horizons  
827 A–C (shaded horizons) indicate correlatable remarkable isotope excursions. Horizons  
828 yielding conodonts after Garzanti *et al.* (1994) are indicated by sequence 1–7; see regional  
829 geology section in text. Abbreviations: G.: Griesbachian; D.: Dienerian; HST: highstand  
830 system tract; TST: transgressive system tract; LST: lowstand system tract; SB: system  
831 boundary.

832 Fig. 2 Photographs and photomicrographs. A: Outcrop of the sample section. Abbreviations:  
833 P: Permian Puchenpra Formation; TB: topmost biocalcarenes; Tc1, Tc2, and Tc3: first,  
834 second, and third carbonate bands of the Tamba Kurkur Formation, respectively; M:

835 Mukut Formation. B: Cross-bedded sandstone in the Puchenpra Formation. C: Contact of  
836 the Puchenpra Formation, topmost biocalcarenes, and first carbonate band of the Tamba  
837 Kurkur Formation. Note the undulated boundary between the Puchenpra Formation and  
838 topmost biocalcarenes. D: Echinoidal wackestone–packstone including quartz grains in  
839 the topmost biocalcarenes. E: Mudstone containing thinly-bedded, alternating layers of  
840 red–purple and black mudstones in the first mudstone interval. F: Small burrow on the  
841 bedding plane of the red–purple mudstone in the first mudstone interval. G: Grading  
842 structure in the red–purple mudstone in the first mudstone interval. The arrow indicates  
843 fining trend of grain size. H: Biomicritic wackestone including small ammonoid (A), thin-  
844 shelled bivalve (B), and ostracoda (O?) in the first carbonate band. I: Biomicritic  
845 packstone–wackestone with thin-shelled bivalves in the second carbonate band. J:  
846 Biomicritic wackestone–packstone, containing fragments of small bivalves in the third  
847 carbonate band. K: Nodular limestone in the third carbonate band. Note the intensely  
848 deformed structure by tectonic–diagenetic brecciation with dissolution seams (dark color).  
849 L: Alternation of marly limestone and mudstone in the Mukut Formation.

850 Fig. 3 Thin-section photomicrograph of mudstones in the Triassic Tamba Kurkur Formation in  
851 the Manang area, central Nepal. a: Sample no. 66 from the first mudstone interval, b:  
852 sample no. 75 from the first mudstone interval, c: sample no. 82 from the second carbonate  
853 band, 4: sample no. 85 from the second carbonate interval. The white triangles indicate  
854 quartz and feldspar grains in mudstone. All photomicrographs were taken under open  
855 polarized light.

856 Fig. 4 Variation of  $\delta^{13}\text{C}_{\text{carb}}-\delta^{18}\text{O}_{\text{carb}}$  for the carbonate in the Triassic topmost biocalcarenes,  
857 Tamba Kurkur, and Mukut Formations in the Manang area, central Nepal.

858 Fig. 5 Variation of major elements against SiO<sub>2</sub> for the mudstones in the Permian Puchenpra,  
859 Triassic topmost biocalcarenes, Tamba Kurkur, and Mukut Formations in the Manang  
860 area, central Nepal.

861 Fig. 6 Chondrite-normalized spider diagram of average REE abundances in the Permian and  
862 Triassic mudstones from the Manang area, central Nepal compared with the PAAS value  
863 (Taylor and McLennan, 1985). Note that all mudstones in the Tamba Kurkur Formation  
864 have a similar source that correlates in composition to felsic igneous rock.

865 Fig. 7 Vertical variation of CIA (Nesbitt and Young, 1982), Al<sub>2</sub>O<sub>3</sub>/Na<sub>2</sub>O ratios, and ICV (Cox  
866 et al., 1995) within the stratigraphic column. Note the visible decrease in CIA and  
867 Al<sub>2</sub>O<sub>3</sub>/Na<sub>2</sub>O values in the second carbonate band (shaded horizon; middle Smithian). The  
868 hatched region in the CIA and Al<sub>2</sub>O<sub>3</sub>/Na<sub>2</sub>O diagram represents the variation of CIA values  
869 in average shales (Taylor and McLennan, 1985). The results of grain size analysis for  
870 mudstone are also shown.

871 Fig. 8 A–CN–K plots (Nesbitt and Young, 1984) for mudstone samples in the Permian  
872 Puchenpra, Triassic topmost biocalcarenes, Tamba Kurkur, and Mukut Formations in the  
873 Manang area, central Nepal. A: Al<sub>2</sub>O<sub>3</sub>, CN: CaO + Na<sub>2</sub>O, K: K<sub>2</sub>O.

874

875 Supplemental Table 1. REE data from the Triassic, Manang area, central Nepal.

876 Supplemental Table 2. Major geochemistry, CIA, Al<sub>2</sub>O<sub>3</sub>/Na<sub>2</sub>O, and ICV data from the  
877 Triassic, Manang area, central Nepal.

878

879

880 Appendix

881 **METHODOLOGY**

882 In this study, analytical methods were primarily used to determine the bulk chemical  
883 composition of major and REEs for 50 mudstone samples. The major element measurements  
884 were obtained by X-ray fluorescence (XRF). For the REE composition, samples were analyzed  
885 by inductively coupled plasma mass spectrometry (ICP-MS). To estimate the geochemical  
886 effects of grain-size variation, the distributions of grain size in mudstones were obtained through  
887 thin-section examinations.

888 **Geochemistry:** To determine the mudstone bulk chemistry, samples were carefully  
889 collected from the mudstone layers of the Puchenpra, topmost biocalcarenes, Tamba Kurkur,  
890 and Mukut Formations. The stratigraphic intervals for sampling ranged from 0.3 to 5 m along the  
891 section. Thin sections were prepared from all mudstone samples. Other parts of the samples were  
892 crushed and ground to a fine powder with a steel and agate mortar, and precautions were taken to  
893 avoid contamination of the samples. The samples were kept at a constant particle size to avoid  
894 the risk of variation in fluorescence intensities. Conventional fused glass disks were favored for  
895 the measurement of major elements because the decomposition of a portion of the sample and  
896 flux can produce a homogeneous glass, thus eliminating the effects of particle size and  
897 mineralogy. The fused glass disks were composed of a 1:2 mixture of powdered dry sample and  
898  $\text{Li}_2\text{B}_4\text{O}_7$  flux, respectively. After being properly mixed, they were heated in Pt crucibles by using  
899 an electric heating system. During fusion, the melt was agitated to produce a homogeneous glass.  
900 For major element analysis, the compositions of all samples were measured by an automatic  
901 XRF spectrometer (Phillips PW2400) at the Faculty of Science, Shinshu University, Japan, using  
902 established procedures and detection limits (Miyake et al., 1996).

903 To clarify the nature of the mudstones, average abundances of REE were examined. Nine  
904 samples from the first mudstone interval, four samples from the second carbonate band, one  
905 sample from the third carbonate band, and three samples from the Mukut Formation were  
906 selected and analyzed. REE analysis was performed following the procedure established by  
907 Roser et al. (2000) with the variation of dissolving fluxed samples in concentrated nitric acid  
908 rather than perchloric acid. Sample powders were dissolved in concentrated HF–HNO<sub>3</sub> by using  
909 Pt crucibles and were then evaporated. Following this step, Na<sub>2</sub>CO<sub>3</sub> alkali flux was added, and  
910 the sample was heated in an electric furnace. The fluxed samples were dissolved in concentrated  
911 HNO<sub>3</sub>, HCl, and ultrapure water. The samples were finally diluted 20,000 times. For calibration,  
912 we used reference values from BHVO-1 (basalt; Hawaii, U.S. Geological Survey [USGS];  
913 Eggins et al., 1997). REE abundances were analyzed by using an inductively coupled plasma  
914 mass spectrometer (Agilent 7500a) at the Faculty of Science, Niigata University, Japan. The  
915 average REE abundances were calculated and are shown as chondrite normalized values in Fig. 6.

916 For the estimation of the geochemical characteristics of the degree of weathering at the  
917 time of deposition, we calculated the chemical index of alteration (CIA) from the major element  
918 geochemistry (Nesbitt and Young, 1982; Fedo et al., 1997). CIA is expressed as  $CIA = \text{molar} [(Al_2O_3)/(Al_2O_3 + CaO^* + Na_2O + K_2O)] \times 100$ , where CaO\* represents CaO in silicate minerals  
919 only. The Al<sub>2</sub>O<sub>3</sub>/Na<sub>2</sub>O ratio is directly calculated from Al<sub>2</sub>O<sub>3</sub> wt% and Na<sub>2</sub>O wt% obtained from  
920 the XRF results. The index of compositional variability (ICV) value, which is expressed as  $ICV = [Fe_2O_3 + K_2O + Na_2O + CaO + MgO + MnO + TiO_2]/Al_2O_3$  (Cox et al., 1995), indicates the  
921 degree of maturity of the fine aluminosiliciclastic material delivered to the sedimentary basin.

924 **Carbon isotope analysis:** Sampling of limestone and marlstone for carbon isotope  
925 measurements was conducted at intervals ranging from 0.3 to 5 m along the sample section.

926 Powders were produced from samples by using a diamond cutter, and hand specimens were  
927 carefully examined to avoid cracks, veins, and weathering features. Fifty-six samples were  
928 analyzed by using an automated carbonate preparation system (carbonate device) connected to  
929 the Thermo-Finnigan MAT253 mass spectrometer at the Kochi Core Center in Kochi University,  
930 Japan. Reproducibility of replicate analyses was better than  $\pm 0.03\%$  for standards and  $\pm 0.2\%$  for  
931 sediment samples of both carbon isotope values. All isotope results are defined as per mill (‰)  
932 deviation vs. PDB and were calibrated to NBS-19 international standards.

933 **Grain-size analysis:** Grain-size analysis was conducted by thin-section examination of seven  
934 selected mudstone samples by using an eye-piece micrometer to measure the dimensions of  
935 quartz and feldspar grains under a standard petrographic microscope of 400× magnification.  
936 Random grains were selected by using a mechanical stage with a grid spacing of 0.2 mm.

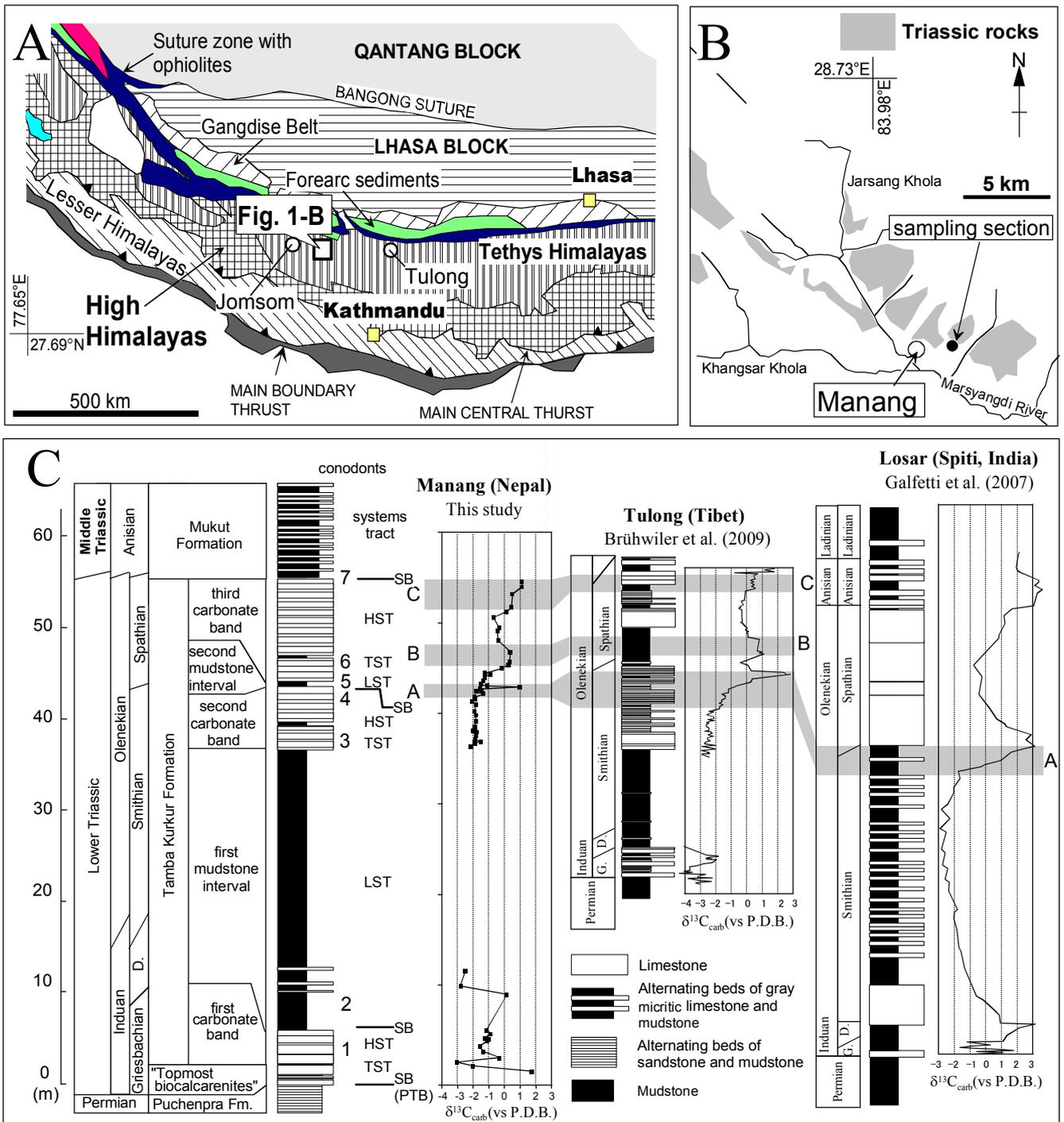


Fig. 1 Yoshida et al.

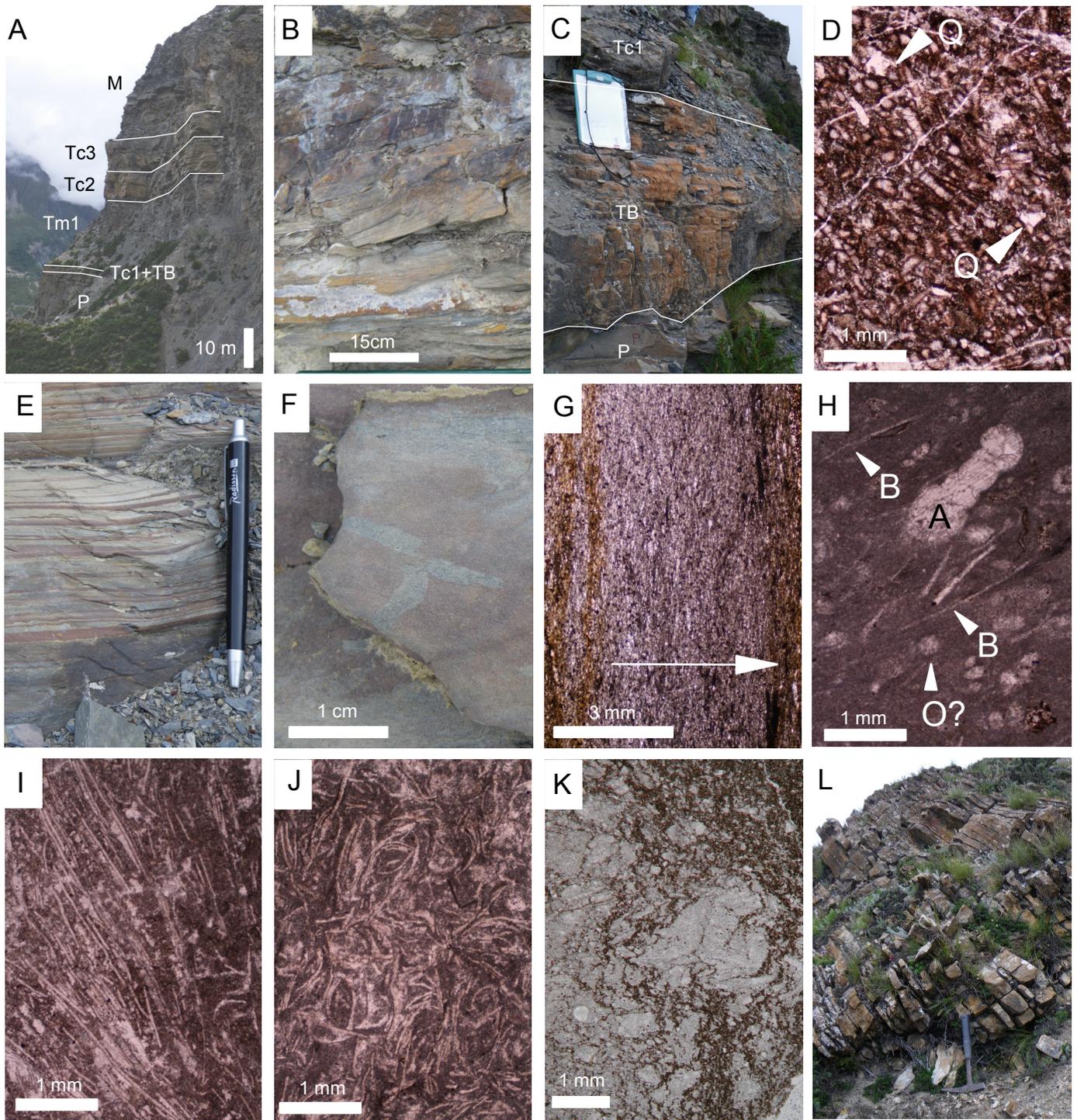


Fig. 2 Yoshida et al.

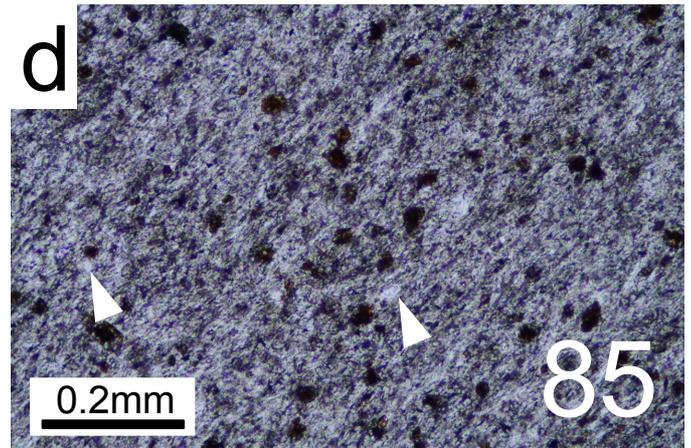
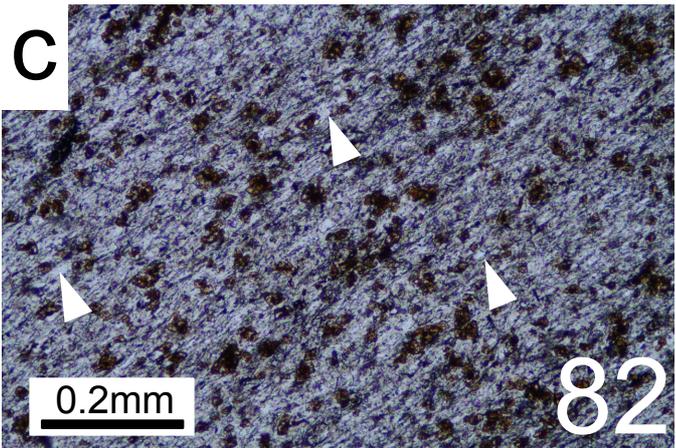
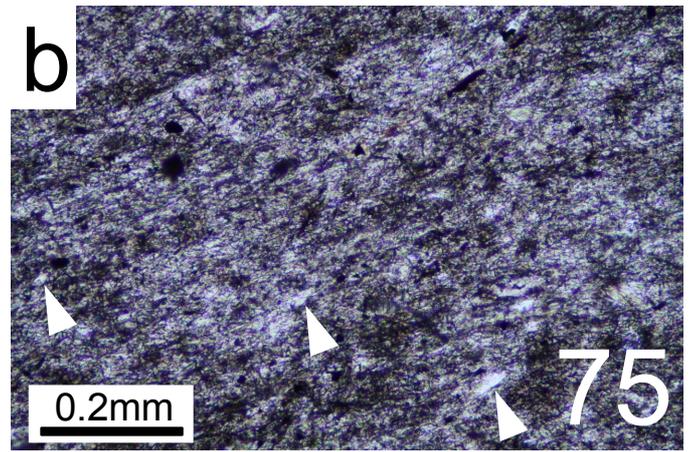
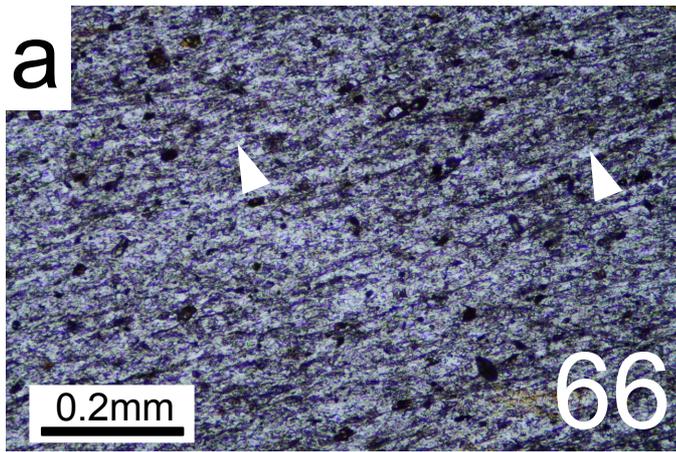


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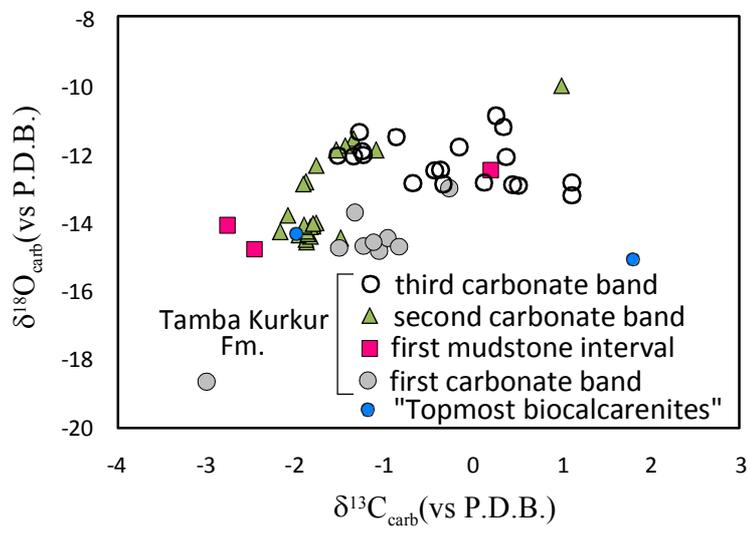


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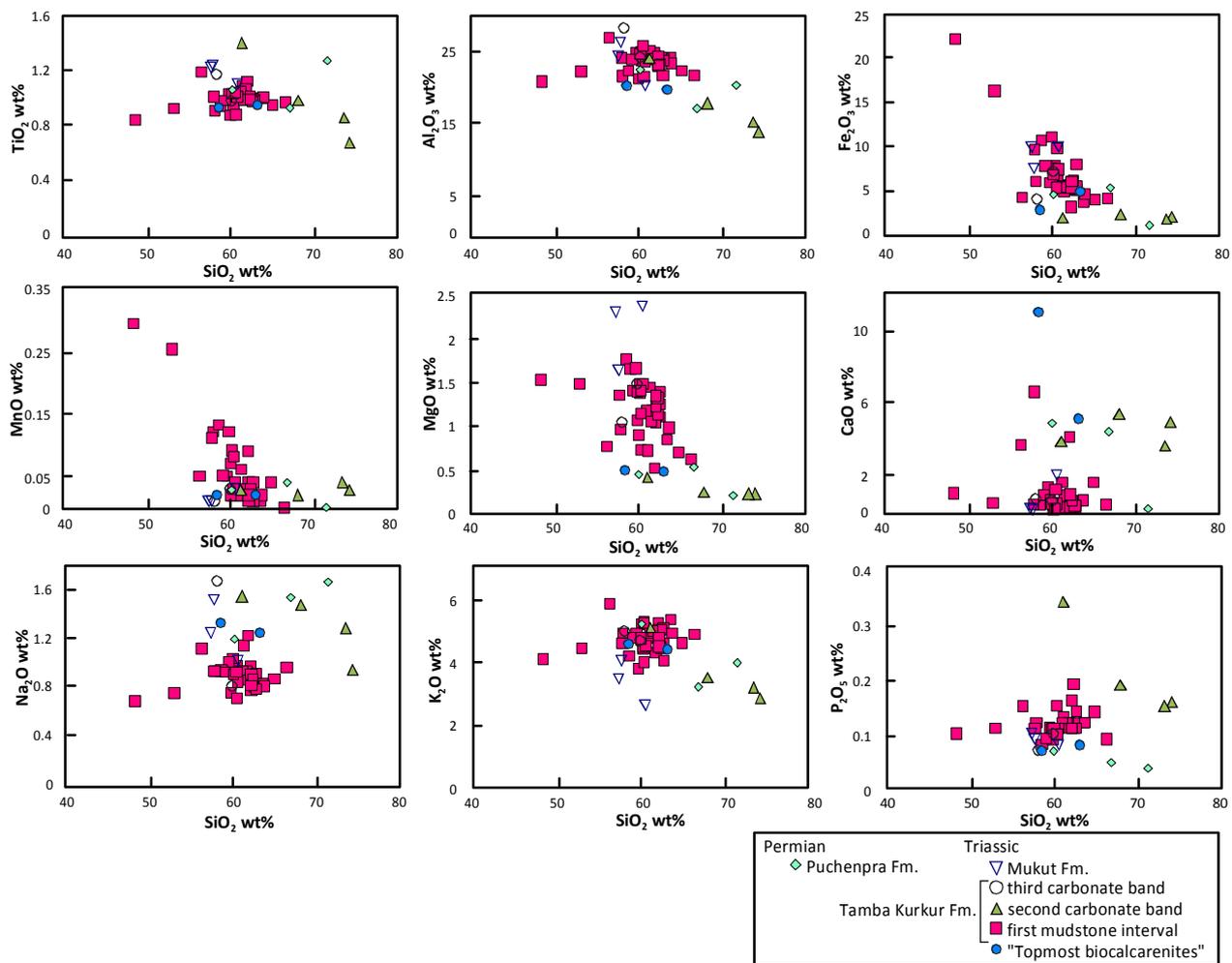


Fig. 5 Yoshida et al.

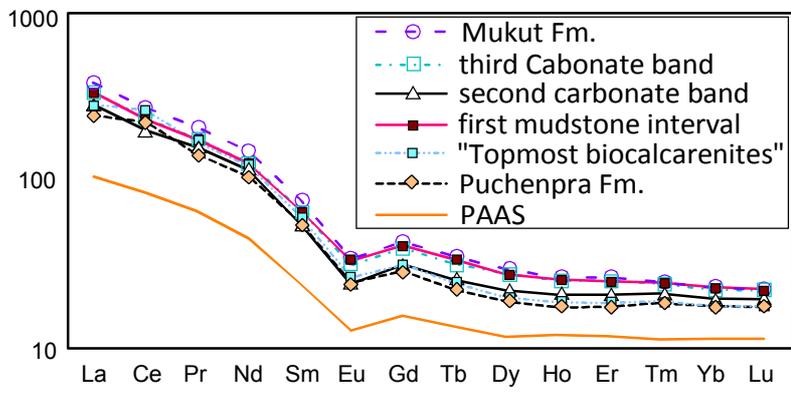


Fig. 6 Yoshida et al.

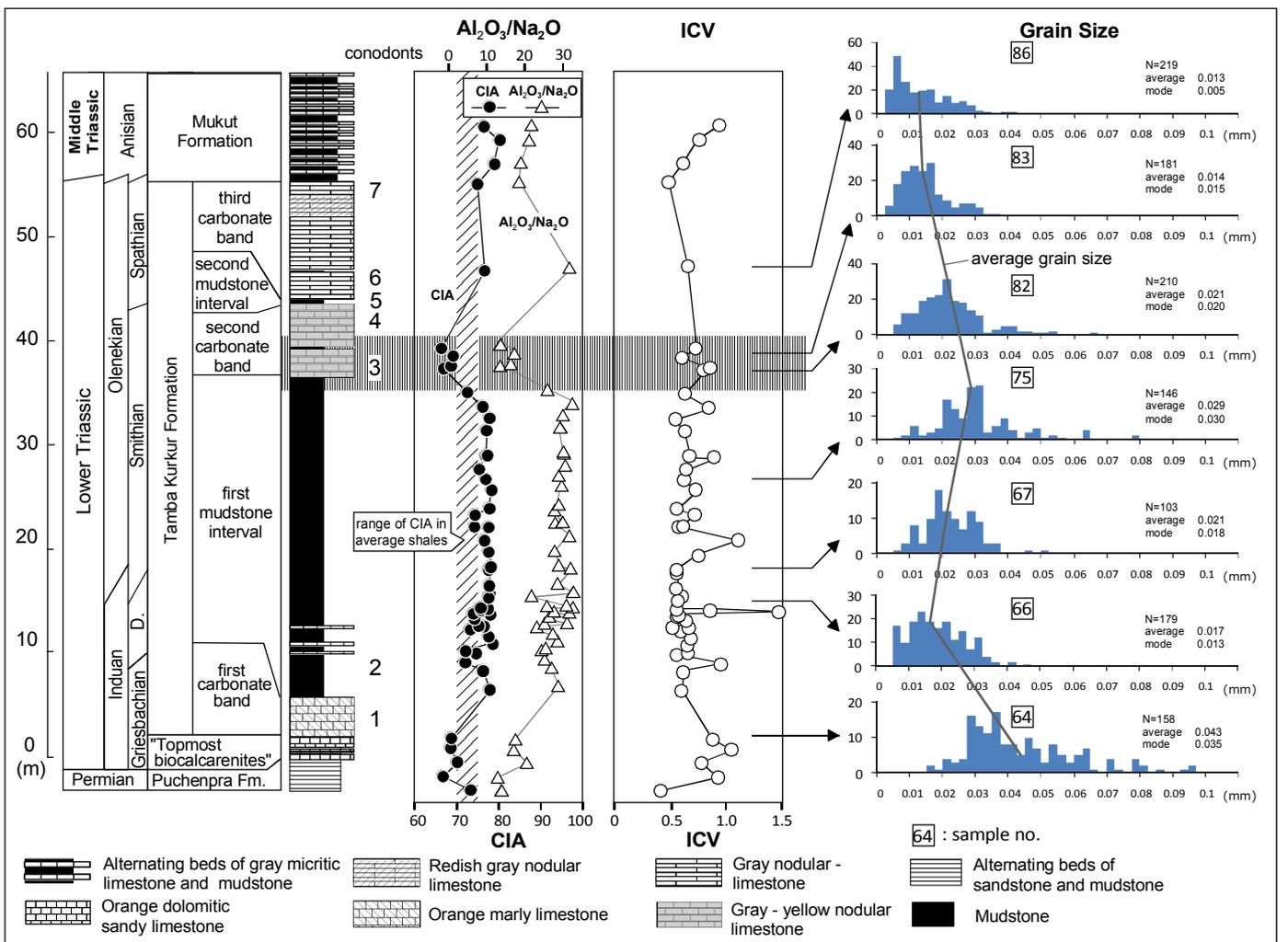


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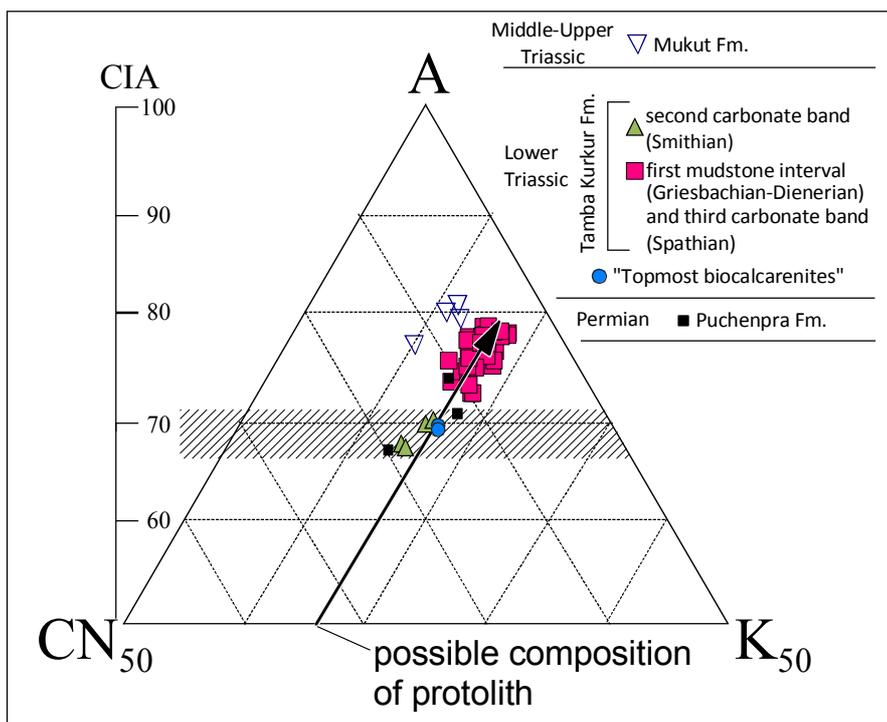


Fig. 8 Yoshida et al.

**Supplemental table 1**

REE data from the Permian-Triassic sediments, Manang area, central Nepal.

	Puchenpra Fm.	Topmost bio-calcarenites	Tamba Kurkur Fm.			Mukut Fm.	PAAS
			First mudstone interval	Second carbonate band	Third carbonate band		
La(N)	243.6	280.6	329.3	283.4	337.2	377.7	103.5
Ce(N)	216.0	258.8	232.1	197.9	232.8	268.9	83.6
Pr(N)	141.9	166.7	177.3	155.6	175.9	204.9	65.0
Nd(N)	106.2	122.0	127.7	112.6	125.7	147.5	45.0
Sm(N)	57.2	61.1	66.9	55.2	66.1	77.0	24.2
Eu(N)	24.8	25.7	32.5	24.2	30.5	33.5	12.6
Gd(N)	28.4	30.7	40.0	31.4	38.4	42.6	15.4
Tb(N)	22.2	24.2	33.6	25.4	31.4	34.7	13.3
Dy(N)	18.8	20.0	28.6	22.2	27.2	29.0	11.5
Ho(N)	17.5	18.6	25.8	20.7	25.2	26.8	11.8
Er(N)	17.3	18.4	25.6	21.0	24.8	25.7	11.6
Tm(N)	18.3	18.6	24.9	21.0	23.8	24.3	11.2
Yb(N)	17.7	17.4	22.8	19.5	22.2	22.3	11.3
Lu(N)	17.6	17.3	22.5	19.6	21.8	21.6	11.3
(La/Yb)N	13.8	16.2	14.5	14.6	15.2	16.9	9.2
(La/Sm)N	4.3	4.6	4.9	5.1	5.1	4.9	4.3
((Gd/Yb)N	1.6	1.8	1.8	1.6	1.7	1.9	1.4

**Note:** (N) indicates chondrite normalized value.

Supplemental table 2

Major bulk rock geochemistry, CIA, Al<sub>2</sub>O<sub>3</sub>/Na<sub>2</sub>O and ICV data from the Permian - Triassic sediments, Manang area, central Nepal.

Sample name	Sampling horizon (m)	System	Formation	Stratigraphic division	SiO <sub>2</sub>	TiO <sub>2</sub>	Al <sub>2</sub> O <sub>3</sub>	Fe <sub>2</sub> O <sub>3</sub>	MnO	MgO	CaO	Na <sub>2</sub> O	K <sub>2</sub> O	P <sub>2</sub> O <sub>5</sub>	Total	CIA	Al <sub>2</sub> O <sub>3</sub> /Na <sub>2</sub> O	ICV
60	-3	Permian	Puchenpra Fm.		71.54	1.28	20.22	0.82	0.00	0.21	0.21	1.66	4.02	0.04	100.00	73.29	12.18	0.41
61	-1.7	Permian	Puchenpra Fm.		66.98	0.92	17.16	5.15	0.04	0.55	4.39	1.53	3.25	0.05	100.00	66.78	11.24	0.92
62	-0.3	Permian	Puchenpra Fm.		60.22	1.05	22.47	4.35	0.03	0.45	4.91	1.19	5.26	0.07	100.00	70.02	18.81	0.77
63	1	Lower Triassic	"Topmost biocalcarenites"		58.58	0.93	20.33	2.72	0.02	0.50	10.95	1.32	4.58	0.07	100.00	68.58	15.37	1.03
64	2	Lower Triassic	"Topmost biocalcarenites"		63.23	0.95	19.59	4.80	0.02	0.49	5.15	1.23	4.46	0.08	100.00	68.78	15.87	0.87
M2	7.2	Lower Triassic	Tamba Kurkur Fm.	First mudstone interval	60.36	0.93	24.86	6.53	0.09	0.89	0.10	0.92	5.22	0.10	100.00	77.76	26.91	0.59
0826M3	9	Lower Triassic	Tamba Kurkur Fm.	First mudstone interval	60.60	1.01	24.43	6.41	0.02	0.72	0.41	0.97	5.28	0.15	100.00	76.08	25.31	0.61
M4	9.8	Lower Triassic	Tamba Kurkur Fm.	First mudstone interval	58.10	0.90	21.51	5.86	0.12	0.96	6.58	0.92	4.94	0.12	100.00	71.98	23.40	0.94
0826M5	10.7	Lower Triassic	Tamba Kurkur Fm.	First mudstone interval	66.62	0.96	21.60	3.89	0.00	0.62	0.37	0.95	4.90	0.09	100.00	74.65	22.69	0.54
M6	10.9	Lower Triassic	Tamba Kurkur Fm.	First mudstone interval	62.28	0.98	22.77	2.93	0.09	0.52	4.09	0.96	5.26	0.12	100.00	72.01	23.74	0.65
0826M7	11.6	Lower Triassic	Tamba Kurkur Fm.	First mudstone interval	60.51	1.02	24.04	7.17	0.02	1.48	0.31	0.90	4.44	0.10	100.00	78.43	26.81	0.64
0826M8	12.3	Lower Triassic	Tamba Kurkur Fm.	First mudstone interval	60.46	1.03	23.60	7.61	0.02	1.37	0.39	0.92	4.50	0.10	100.00	77.50	25.53	0.67
65	13	Lower Triassic	Tamba Kurkur Fm.	First mudstone interval	61.44	1.08	24.29	4.73	0.06	0.72	1.61	1.13	4.80	0.13	100.00	73.13	21.45	0.58
MA-1(blk)	13.3	Lower Triassic	Tamba Kurkur Fm.	First mudstone interval	60.18	1.01	24.01	7.37	0.07	1.06	0.47	1.02	4.71	0.09	100.00	76.42	23.48	0.65
M9	13.35	Lower Triassic	Tamba Kurkur Fm.	First mudstone interval	63.77	0.99	23.98	3.52	0.01	0.84	0.59	0.82	5.37	0.12	100.00	75.15	29.30	0.51
MA-2(blk)	14	Lower Triassic	Tamba Kurkur Fm.	First mudstone interval	59.74	1.02	24.65	5.76	0.05	1.41	1.33	1.00	4.93	0.11	100.00	74.09	24.69	0.63
0826M10	14.4	Lower Triassic	Tamba Kurkur Fm.	First mudstone interval	62.91	0.99	23.92	4.89	0.01	1.10	0.14	0.80	5.11	0.12	100.00	77.83	29.97	0.54
66	14.5	Lower Triassic	Tamba Kurkur Fm.	First mudstone interval	65.12	0.94	22.20	3.77	0.04	0.70	1.62	0.86	4.62	0.14	100.00	73.92	25.86	0.56
0826M11A	14.9	Lower Triassic	Tamba Kurkur Fm.	First mudstone interval	48.54	0.83	20.73	22.18	0.30	1.53	1.01	0.67	4.11	0.10	100.00	75.68	30.86	1.46
MA-3(blk)	15	Lower Triassic	Tamba Kurkur Fm.	First mudstone interval	58.79	0.94	22.23	10.59	0.13	1.76	0.35	0.93	4.20	0.08	100.00	77.35	24.00	0.84
0826M11B	15.1	Lower Triassic	Tamba Kurkur Fm.	First mudstone interval	63.95	1.00	23.19	4.39	0.02	0.98	0.62	0.80	4.93	0.12	100.00	75.59	29.13	0.55
0826M12	16.4	Lower Triassic	Tamba Kurkur Fm.	First mudstone interval	62.38	0.97	23.51	5.85	0.04	1.03	0.24	0.76	5.05	0.16	100.00	77.65	30.93	0.59
67	16	Lower Triassic	Tamba Kurkur Fm.	First mudstone interval	62.00	1.11	24.26	5.40	0.02	1.17	0.37	1.21	4.32	0.12	100.00	77.44	19.98	0.56
68	17.2	Lower Triassic	Tamba Kurkur Fm.	First mudstone interval	61.87	1.05	24.68	5.04	0.02	1.05	0.29	0.92	4.95	0.12	100.00	77.61	26.82	0.54
M14	18.7	Lower Triassic	Tamba Kurkur Fm.	First mudstone interval	62.70	0.97	24.03	4.95	0.02	1.13	0.26	0.79	5.04	0.11	100.00	77.51	30.24	0.55
69	19	Lower Triassic	Tamba Kurkur Fm.	First mudstone interval	61.30	1.04	24.94	5.27	0.02	1.17	0.26	0.92	4.95	0.12	100.00	77.94	27.12	0.55
70	20.4	Lower Triassic	Tamba Kurkur Fm.	First mudstone interval	58.00	1.00	23.98	9.52	0.11	1.35	0.41	0.92	4.60	0.11	100.00	77.44	26.04	0.74
0826M16	20.9	Lower Triassic	Tamba Kurkur Fm.	First mudstone interval	53.20	0.91	22.14	16.21	0.25	1.48	0.49	0.74	4.47	0.11	100.00	76.84	29.92	1.10
71	22.2	Lower Triassic	Tamba Kurkur Fm.	First mudstone interval	62.94	0.98	23.64	5.25	0.01	1.24	0.33	0.90	4.59	0.11	100.00	77.70	26.25	0.56
0826M17	22.25	Lower Triassic	Tamba Kurkur Fm.	First mudstone interval	61.62	0.98	23.85	5.22	0.03	1.44	0.89	0.85	5.01	0.11	100.00	74.49	28.14	0.60
M18	23.4	Lower Triassic	Tamba Kurkur Fm.	First mudstone interval	59.29	0.97	23.74	7.60	0.05	1.65	0.90	0.91	4.79	0.09	100.00	74.56	26.10	0.71
72	24	Lower Triassic	Tamba Kurkur Fm.	First mudstone interval	62.37	1.00	24.22	5.15	0.01	1.23	0.25	0.89	4.77	0.11	100.00	78.01	27.16	0.55
73	25.8	Lower Triassic	Tamba Kurkur Fm.	First mudstone interval	62.94	0.97	21.52	7.85	0.04	1.39	0.31	0.77	4.07	0.14	100.00	78.48	27.92	0.71
74	26.8	Lower Triassic	Tamba Kurkur Fm.	First mudstone interval	62.59	0.96	23.06	5.92	0.03	1.32	0.60	0.85	4.47	0.19	100.00	77.05	27.07	0.61
75	27.8	Lower Triassic	Tamba Kurkur Fm.	First mudstone interval	62.36	0.98	23.03	5.90	0.02	1.34	0.95	0.80	4.51	0.11	100.00	75.40	28.86	0.63
76	29	Lower Triassic	Tamba Kurkur Fm.	First mudstone interval	59.98	0.86	21.14	10.93	0.12	1.65	0.66	0.74	3.81	0.11	100.00	77.12	28.53	0.88
77	29.14	Lower Triassic	Tamba Kurkur Fm.	First mudstone interval	60.80	1.00	23.54	7.23	0.04	1.48	0.45	0.83	4.52	0.11	100.00	77.57	28.30	0.66
78	31.5	Lower Triassic	Tamba Kurkur Fm.	First mudstone interval	60.22	1.02	24.50	6.63	0.02	1.39	0.45	0.89	4.77	0.11	100.00	77.36	27.50	0.62
79	32.7	Lower Triassic	Tamba Kurkur Fm.	First mudstone interval	60.56	1.03	25.66	5.14	0.03	1.14	0.16	0.91	5.25	0.10	100.00	78.01	28.18	0.53
80	33.8	Lower Triassic	Tamba Kurkur Fm.	First mudstone interval	60.63	0.87	21.36	9.61	0.08	1.41	1.23	0.70	4.00	0.10	100.00	76.32	30.62	0.83
81	35.2	Lower Triassic	Tamba Kurkur Fm.	First mudstone interval	56.48	1.18	26.74	4.01	0.05	0.76	3.65	1.10	5.87	0.15	100.00	72.81	24.23	0.62
82	37.5	Lower Triassic	Tamba Kurkur Fm.	Second carbonate band	68.19	0.98	17.74	2.18	0.02	0.25	5.42	1.47	3.55	0.19	100.00	67.10	12.03	0.78
83	37.7	Lower Triassic	Tamba Kurkur Fm.	Second carbonate band	74.30	0.68	13.85	1.97	0.03	0.22	4.96	0.94	2.89	0.16	100.00	69.01	14.76	0.84
84	38.7	Lower Triassic	Tamba Kurkur Fm.	Second carbonate band	61.28	1.40	24.12	1.78	0.03	0.42	3.93	1.54	5.16	0.34	100.00	69.39	15.69	0.59
85	39.9	Lower Triassic	Tamba Kurkur Fm.	Second carbonate band	73.53	0.86	15.36	1.71	0.04	0.22	3.66	1.27	3.20	0.15	100.00	66.75	12.07	0.71
86	47	Lower Triassic	Tamba Kurkur Fm.	Third carbonate band	60.12	0.97	24.12	7.07	0.03	1.48	0.60	0.80	4.71	0.10	100.00	76.87	30.14	0.65
87	55.3	Upper Triassic	Mukut Fm.		58.29	1.16	28.18	3.85	0.01	1.04	0.74	1.66	5.00	0.07	100.00	75.15	16.98	0.48
88	57.4	Upper Triassic	Mukut Fm.		57.88	1.23	26.12	7.34	0.01	1.64	0.11	1.51	4.07	0.09	100.00	79.19	17.33	0.61
89	59.6	Upper Triassic	Mukut Fm.		57.55	1.22	24.20	9.75	0.01	2.30	0.17	1.24	3.47	0.10	100.00	80.50	19.56	0.75
90	60.9	Upper Triassic	Mukut Fm.		60.76	1.10	20.21	9.82	0.03	2.36	2.00	1.00	2.64	0.08	100.00	76.65	20.13	0.94